

**DEPARTMENT OF PHYSICS  
INDIAN INSTITUTE OF TECHNOLOGY, MADRAS**

PH587 GR & Cosmology

Assignment 3

15.3.2009 (due: 23.3.2009)

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1. Consider the spherical coordinate system for  $\mathbb{R}^3$ .
  - (a) After obtaining the metric in this coordinate system, compute all the components of the Christoffel symbol for the metric.
  - (b) Compute the covariant derivative of a vector field,  $v^a$ , in this coordinate system and hence obtain its divergence,  $v^a{}_{;a}$ . Compare your result with known results for the divergence. *Hint:* You will need to suitably rescalings to obtain a match. Why?
  - (c) Also obtain the expression for the scalar Laplacian and compare with standard formulae.

2. Prove the following identities:

- (a)  $\Gamma_{ab}^b = \frac{1}{2} \partial_a \ln g$ .
- (b)  $g^{ab} \Gamma_{ab}^c = -(g^{cd} \sqrt{g})_{,b} / \sqrt{g}$ .
- (c) For an antisymmetric tensor  $F^{ab}$ ,  $F^{ab}{}_{;b} = (\sqrt{g} F^{ab})_{,b} / \sqrt{g}$ .
- (d)  $g^{ab} g_{bc,d} = -g^{ab}{}_{,d} g_{bc}$ .
- (e)  $g^{ab}{}_{,c} = -\Gamma_{cd}^a g^{db} - \Gamma_{cd}^b g^{da}$ .

3. **Riemann Normal Coordinates:** Consider a Riemannian manifold with metric  $g_{ab}(x)$ . Let  $P$  be a point with coordinate  $x_0^a$  on the manifold. Under a general coordinate transformation,  $x'^a(x)$ , the metric tensor transforms as

$$g'_{a_1 a_2}(x') = \frac{\partial x^{b_1}}{\partial x'^{a_1}} \frac{\partial x^{b_2}}{\partial x'^{a_2}} g_{b_1 b_2}(x) .$$

Consider the Taylor expansion about the point  $P$

$$\begin{aligned} \delta x^b(x') = \frac{\partial x^b}{\partial x'^{a_1}} \Big|_P \delta x'^{a_1} + \frac{1}{2} \frac{\partial^2 x^b}{\partial x'^{a_1} \partial x'^{a_2}} \Big|_P \delta x'^{a_1} \delta x'^{a_2} \\ + \frac{1}{3!} \frac{\partial^3 x^b}{\partial x'^{a_1} \partial x'^{a_2} \partial x'^{a_3}} \Big|_P \delta x'^{a_1} \delta x'^{a_2} \delta x'^{a_3} + \dots \end{aligned}$$

where  $\delta x^b(x') = x^b(x') - x_0^b$  and  $\delta x'^b = x'^b - x_0'^b$ . We will now consider a sequence of three coordinate transformations  $x \rightarrow x' \rightarrow x'' \rightarrow x'''$  where we will attempt to bring the metric in the neighbourhood of  $P$  as close to  $\delta_{ab}$  as possible.

(a) **Step 1: Set  $g'_{ab}(P) = \delta_{ab}$ :** Let  $A$  be a constant matrix satisfying

$$\delta_{a_1 a_2} = A_{a_1}^{b_1} A_{a_2}^{b_2} g_{b_1 b_2}(P) .$$

Show that under the (inverse) coordinate transformation

$$\delta x^b(x') = A_{a_1}^{b_1} \delta x'^{a_1} ,$$

$$g'_{a_1 a_2}(P) = \delta_{a_1 a_2} .$$

(b) **Step 2: Set  $g''_{ab}(x'') = \delta_{ab} + \mathcal{O}(\delta x''^2)$ :** Consider the (inverse) coordinate transformation

$$\delta x'^b(x'') = \delta x''^b + \frac{1}{2} B_{a_1 a_2}^b \delta x''^{a_1} \delta x''^{a_2} ,$$

with the constant hypermatrix satisfying  $B_{a_1 a_2}^b = B_{a_2 a_1}^b$ . Show that

$$g''_{a_1 a_2}(x'') = \delta_{a_1 a_2} + (g'_{a_1 a_2, c}(P) + B_{a_1 c}^b \delta_{b a_2} + B_{a_2 c}^b \delta_{b a_1}) \delta x''^c + \mathcal{O}(\delta x''^2) .$$

In other words,

$$g''_{a_1 a_2, c}(P) = (g'_{a_1 a_2, c}(P) + B_{a_1 c}^b \delta_{b a_2} + B_{a_2 c}^b \delta_{b a_1})$$

Show that  $g''_{a_1 a_2, c}(P)$  vanishes if we choose  $B_{ac}^b = -\Gamma_{ac}^b(P)$ .

(c) **Step 3:** Consider the (inverse) coordinate transformation

$$\delta x''^b(x''') = \delta x'''^b + \frac{1}{6} C_{a_1 a_2 a_3}^b \delta x'''^{a_1} \delta x'''^{a_2} \delta x'''^{a_3} ,$$

with the constant hypermatrix  $C_{a_1 a_2 a_3}^b$  being totally symmetric in the indices  $(a_1, a_2, a_3)$ . Show that

$$g'''_{a_1 a_2}(x''') = \delta_{a_1 a_2} + \frac{1}{2} \left( g''_{a_1 a_2, c_1 c_2}(P) + C_{a_1 c_1 c_2}^b \delta_{b a_2} + C_{a_2 c_1 c_2}^b \delta_{b a_1} \right) \delta x'''^{c_1} \delta x'''^{c_2} + \mathcal{O}(\delta x'''^3) . \quad (1)$$

In other words

$$g'''_{a_1 a_2, c_1 c_2}(P) = (g''_{a_1 a_2, c_1 c_2}(P) + C_{a_1 c_1 c_2}^b \delta_{b a_2} + C_{a_2 c_1 c_2}^b \delta_{b a_1})$$

(d) Show that the combination

$$R''_{a_1 c_1 a_2 c_2} \equiv \frac{1}{2} \left[ -g''_{a_1 a_2, c_1 c_2}(P) + g''_{c_1 a_2, a_1 c_2}(P) + g''_{a_1 c_2, c_1 a_2}(P) - g''_{c_1 c_2, a_1 a_2}(P) \right]$$

is left unaffected by the third coordinate transformation i.e.,  $R'''_{a_1 c_1 a_2 c_2} = R''_{a_1 c_1 a_2 c_2}$ . This shows that there exists **no** coordinate transformation which can set this combination to zero. This combination is called the **Riemann curvature tensor**.

- (e) The tensorial properties of the Riemann tensor are not quite evident as it is computed in a special coordinate system where  $g''_{ab,c}(P) = 0$ . However, verify that the following holds in this coordinate system (and hence holds in all coordinate systems)

$$\begin{aligned} R_{a_1 c_1 a_2 c_2} &= -R_{c_1 a_1 a_2 c_2} = -R_{a_1 c_1 c_2 a_2} = +R_{a_2 c_2 a_1 c_1} , \\ R_{a_1 c_1 a_2 c_2} + R_{a_1 a_2 c_2 c_1} + R_{a_1 c_2 c_1 a_2} &= 0 . \end{aligned}$$

Thus in the neighbourhood of a point, the best one can do is to find a coordinate transformation such that the metric takes the following form:

$$g_{ab}(x) = \delta_{ab} - \frac{1}{3} R_{abcd} \delta x^c \delta x^d + \mathcal{O}(\delta x^3) .$$

Such coordinates are called the **Riemann normal coordinates** in Riemannian geometry. In Lorentzian geometry that is appropriate to general relativity, the analogous statement is that the metric can be brought to the form

$$g_{\mu\nu}(x) = \eta_{\mu\nu} - \frac{1}{3} R_{\mu\rho\nu\sigma} \delta x^\rho \delta x^\sigma + \mathcal{O}(\delta x^3) .$$

Such coordinates are also called **locally inertial coordinates**.

4. A four-dimensional Lorentzian manifold has coordinates  $(t, x, y, z)$  and metric

$$ds^2 = -(1 + 2\phi)dt^2 + (1 - 2\phi)(dx^2 + dy^2 + dz^2) ,$$

where  $|\phi(t, x, y, z)| \ll 1$  everywhere.

- (a) Compute all components of the Christoffel connection to first order in  $\phi$  and its derivatives.
- (b) At any point  $P$  with coordinates  $(t_0, x_0, y_0, z_0)$ , find a coordinate transformation to a locally inertial coordinate system, to first order in  $\phi$ .
- (c) At what rate does this frame accelerate with respect to the original coordinates, again to first order in  $\phi$ . *Hint:* Compare this metric with the one we obtained for a uniformly accelerating observer in class.